Nonlinear Slow Shear Alfven Waves in Electron-Positron-Ion Plasma Including Full Ion Dynamics

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Abstract—Propagation of arbitrary amplitude nonlinear Alfven waves has been investigated in low but finite $\beta$ electron-positron-ion plasma including full ion dynamics. Using Sagdeev pseudopotential method an energy integral equation has been derived. The Sagdeev potential has been calculated for different plasma parameters and it has been shown that inclusion of ion parallel motion along the magnetic field changes the nature of slow shear Alfven wave solitons from dip type to hump type. The effects of positron concentration, plasma-$\beta$ and obliqueness of the wave propagation on the solitary wave structure have also been examined.

Keywords—Alfven waves, Sagdeev potential, Solitary waves.

I. INTRODUCTION

The study of nonlinear Alfven waves in electron-positron-ion plasmas has become of interest in plasma research because Alfven wave activity can be observed in a wide variety of astrophysical plasmas containing electrons, positrons and ions. Large amplitude Alfven wave research attracted renewed attention [1], [2] because of the observation of the Alfven wave activity in the auroral regions by the Freja satellite [3], [4]. The nonlinear solitary structure of kinetic Alfven wave was first analytically studied by [5] in electron-ion plasma. They showed that density humps can be associated with nonlinear Alfven wave dynamics. Thereafter a number of authors have reconsidered the problem by including various factors influencing the solitary wave structure [6]-[10]. Nonlinear dynamics of slow shear Alfven wave was also studied by [11] in low-$\beta$ plasma. It was found that density dips can be associated with this wave in the nonlinear regions. The importance of slow shear Alfven wave was emphasized long ago by Lysak and Carlson [12] in connection with the magnetospheric-ionospheric coupling. All these works are restricted to simple electron-ion plasmas. Most of the astrophysical plasmas contain electrons, ions as well as positrons. The properties of wave motion in e-p-i plasmas are expected to be different from those in two components electron-ion or electron-positron plasmas [13], [14]. For example, kinetic Alfven wave in electron-positron-ion plasma can form density dips while in electron-plasma they form density humps [5]. Recently nonlinear Alfven waves in electron-positron-ion plasmas have been studied by a few authors [15], [16].

Saleem and Mahmood [17] have studied nonlinear kinetic Alfven wave in electron-positron-ion plasma ignoring ion parallel motion. On the other hand, [18] considered solitary kinetic Alfven waves in electron-positron-ion plasma including parallel ion motion and current. Mahmood and Saleem [19] investigated nonlinear solitary structures of arbitrary amplitude slow shear Alfven wave in electron-positron-ion plasma ignoring ion parallel motion. They have shown that electron density dips of SSAW are formed in super Alfvenic region. As ion inertia plays an important role in Alfven wave dynamics due to ion polarization drift, it would be interesting to study nonlinear slow shear Alfven wave in electron-positron-ion plasma including full ion dynamics.

In this paper we present a detailed study of nonlinear slow shear Alfven wave in electron-positron-ion plasma with $\beta \ll (m_e/m_i) <<1$ in which ion inertia and current along the magnetic field are taken into consideration. We show that unlike the previous work [19] which ignored ion parallel motion and predicted dip type density soliton, here hump type solitons for slow shear Alfven wave may be excited.

The paper is organized as follows: Basic equations are presented in Section II. In Section III, we derive the energy integral equation and Sagdeev Potential. In Section IV, we present the numerical analysis graphically and discuss the results.

II. BASIC EQUATION

We consider cold homogeneous electron-positron-ion plasma in presence of a stationary ambient magnetic field along z-axis i.e., $B_0 = B_0 \hat{z}$. The equations governing the dynamics of the nonlinear slow shear Alfven waves in x-z plane in a low-$\beta(=\pi n_i T / B_0^2)$ electron-positron-ion plasma in the low frequency limit are:

For electrons:

$$\frac{\partial n_e}{\partial t} + \frac{\partial (n_e v_e)}{\partial z} = 0$$

$$\frac{\partial v_e}{\partial t} + v_e \frac{\partial v_e}{\partial z} = \frac{\beta}{2Q} \frac{\partial \psi}{\partial z}$$

where $\psi$ denotes the electromagnetic potential associated with the parallel electric field $E_p$, $Q=m_e/m_i$ is the electron to ion mass ratio and $V_e$ is the parallel component of electron velocity with respect to the external magnetic field. In writing
these equations we have assumed that electron inertia dominates over electron pressure.

For positrons:

$$\frac{\partial n_p}{\partial t} + \frac{\partial (n_p v_{pz})}{\partial z} = 0$$

$$\frac{\partial v_{pz}}{\partial t} + v_{pz} \frac{\partial v_{pz}}{\partial z} = -\frac{\beta}{2} \frac{\partial \psi}{\partial z}$$

(3)

(4)

Here, $Q=m_p/m_i$ is the positron to ion mass ratio and $v_{pz}$ is the parallel component positron velocity.

For ions:

$$\frac{\partial n_i}{\partial t} + \frac{\partial (n_i v_{ix})}{\partial z} + \frac{\partial (n_i v_{iz})}{\partial x} = 0$$

$$\frac{\partial v_{ix}}{\partial t} + v_{ix} \frac{\partial v_{ix}}{\partial z} + v_{iz} \frac{\partial v_{iz}}{\partial x} = -\frac{\beta}{2} \frac{\partial \psi}{\partial x}$$

$$v_{ix} = -\frac{\beta}{2} \frac{\partial \phi}{\partial x}$$

(5)

(6)

(7)

Here $v_{ix}$ and $v_{ix}$ are respectively the perpendicular and parallel component of the velocity of ions with respect to the external magnetic field; $\phi$ is the electrostatic potential associated with the perpendicular component $E_x$ of the electric field. The above equations have been written in dimensionless form by normalizing number density ($n_a$) by unperturbed number density ($n_{a0}$), time ($t$) by $1/\Omega$, velocity by Alfvén speed ($v_A$) and potentials ($\phi$) by ($k_B T_e/e$). Here $\Omega$, $k_B$ are the ion gyro-frequency and Boltzmann constant respectively.

This ion velocity is taken to be solely due to polarization drift which is necessary for an electron-positron-ion plasma to maintain the quasineutrality condition. It may be noted here that the electron and positron polarization drifts can be ignored for such plasmas. The expression for the parallel current density is

$$j_z = n_e v_{ez} + n_p v_{pz} - n_i v_{iz}$$

Ampere’s law for $j_z$ gives

$$\nabla_x \cdot \left( \frac{\partial (\phi - \psi)}{\partial z} \right) = \frac{\partial j_z}{\partial t}$$

(9)

The quasineutrality condition is prescribed by

$$n_p + n_i = n_e$$

(10)

III. DERIVATION OF SAGDEEV POTENTIAL EQUATION

We seek stationary localized planar solution of the nonlinear equations given by (1)-(10). So we transform these equations in a moving frame $\eta$ defined by

$$\eta = k_1 x + k_2 z - Ut$$

(11)

Here $U$ is the velocity of the nonlinear structure in the moving frame; $k_1$ and $k_2$ denote the direction cosines in the perpendicular direction i.e., $x$ direction and the parallel direction i.e., the $z$-direction respectively such that

$$k_1^2 + k_2^2 = 1.$$  

(12)

Now the equation for electron continuity (1) and the equation for positron continuity (3) with the use of the transformation given by (11) and appropriate boundary conditions i.e.,

$$|\eta| \to \infty; \quad n_e, n_p \to 1; \quad v_{ez}, v_{pz} \to 0,$$

yield

$$v_{ez} = \frac{U}{k_1}(1 - \frac{1}{n_e})$$

$$v_{pz} = \frac{U}{k_2}(1 - \frac{1}{n_p})$$

(13)

Here $n_e$ and $n_p$ are the electron density and positron density normalized to $n_{e0}$ and $n_{p0}$ respectively. Using (12) and (13) in the equations for electron and positron momentum balance equations (2) and (4) we obtain

$$\frac{U^2}{k_1^2} \frac{1}{n_e^2} = \frac{k^2}{Q}$$

$$\frac{U^2}{k_2^2} \frac{1}{n_p^2} = \frac{k^2}{Q}$$

(14)

(15)

From the equation for ion continuity (5) we get

$$k_1 v_{ix} + k_2 v_{iz} = U(1 - \frac{1}{n_i})$$

(16)

where $n_i = n_i/n_{i0}$; $n_{i0}$ being the unperturbed ion density. In deriving (16), the transformation given by (11) and boundary conditions

$$n_i \to 1, \quad v_{ix}, v_{iz} \to 0 \text{ as } |\eta| \to \infty,$$

has been used.

Combining (14) and (15) one obtains

$$n_p = \frac{1}{\sqrt{2 - \frac{1}{n_{p0}^2}}}$$

(17)

Using the transformation (11) and the relation (16) we obtain the following equation from the equation for ion momentum (6):
\[ \frac{\partial \psi}{\partial \eta} = \frac{\beta k_i n_i}{2U} \frac{\partial \psi}{\partial \eta} \]  
\hspace{2cm} (18)

Now the quasi-neutrality condition (10) expressed in terms of the normalized number densities of the species takes the form
\[ n_i = \frac{n_i - p n_i}{1 - p} \]  
\hspace{2cm} (19)

where \( p = \frac{n_{i0}}{n_{e0}} \) is the ratio of the unperturbed positron and electron densities. Using (17) and (19) we obtain from (18)
\[ \frac{\partial \psi}{\partial \eta} = \frac{p}{2U(1 - p)} \]  
\hspace{2cm} (20)

On combining (14) and (20) one gets
\[ v_i = \frac{UQ}{k_i(1 - p)n_i} + \frac{pUQ}{k_i(1 - p)V} \left( \frac{1}{n_i^2} - \frac{1}{n_i^2} \right) \]  
\hspace{2cm} (21)

Now from (16) we obtain the transverse component velocity as:
\[ v_i = U[1 - \left( \frac{1}{1 - p} \right) \frac{1}{n_i^2} \frac{1}{n_i^2} - p] \]  
\hspace{2cm} (22)

Again (7) under the transformation (11) turns out to be
\[ v_i = \frac{1}{2}\langle Uk_i \beta \frac{\partial^2 \phi}{\partial \eta^2} \rangle \]  
\hspace{2cm} (23)

Comparing (23) and (22) one readily obtains
\[ \frac{\partial^2 \phi}{\partial \eta^2} = \frac{2}{Qk_i \left[ 1 - (1 - p) \left( \frac{2}{n_i^2} - 1 - p \right) \right]} \]  
\hspace{2cm} (24)

Equation (9) in the moving frame can be written as
\[ k_i^2 k_i \left( \frac{\partial^2 \phi}{\partial \eta^2} - \frac{\partial \psi}{\partial \eta} \right) = -U \]  
\hspace{2cm} (26)

Finally using (8), (13), (21), (24) and (25) we obtain from (26), the desired “energy law” in the following form
\[ \frac{1}{2} \left( \frac{dn_i}{d\eta} \right)^2 + V(n_i) = 0 \]  
\hspace{2cm} (27)

where \( V(n_i) \) is the effective Sagdeev potential given by
\[ V(n_i) = \int [\frac{1}{1 - \frac{k_i}{U}\frac{n_i}{n_{e0}} 2n_{e0}^2 - 1 - \frac{1}{2n_{e0}^2}}] \]  
\hspace{2cm} (28)

Equation (28) contains several constants \( A, B, C \) and \( D \) which depend on different plasma parameters such as positron concentration, plasma \( \beta \) and obliqueness of wave propagation. These constants are given by
\[ A = \frac{1}{n_i} \frac{1}{2} \left( 1 + \frac{1}{n_i^2} \right) \]  
\hspace{2cm} (29)

\[ B = \frac{1}{n_i^2} \frac{1}{2} \left( 2n_{e0}^2 - 1 \right) \]  
\hspace{2cm} (30)

\[ C = \frac{1}{n_i^3} \frac{1}{2n_{e0}^2} \]  
\hspace{2cm} (31)

\[ D = \frac{1}{n_i} \frac{1}{2} n_{e0}^2 \]  
\hspace{2cm} (32)
\[ B = \sin^{-1} \left( \frac{1}{2n_e} \right) - \frac{1}{n_e} \sqrt{2 - \frac{1}{n_e}} \]
\[ \frac{\pi}{4} = \frac{1}{2} \left( \frac{1}{n_e} - 1 \right) + \frac{p}{2} \left( \frac{1}{n_e^2} - 1 \right) + \frac{p}{2} \sqrt{2 - \frac{1}{n_e^2}} \]
\[ C = \frac{1}{2} \left( \frac{1}{n_e^2} - 1 \right) \]
\[ - (1 - p) \left( \frac{dn_e}{n_e^2 - 1} \right) \]

and

\[ D = \frac{1}{3} \left( \frac{2 - 1}{n_e^2} \right) - \frac{5}{6} + \frac{1}{2n_e^2} \]

### IV. RESULTS AND DISCUSSION

Equation (27) may be identified as the energy equation of an oscillating quasi-particle of unit mass, velocity \( \frac{dn_e}{dn} \) and position \( n_e \) in potential \( V(n_e) \). For solitary wave solutions, particle motion represented by (27) must be confined between two points \( n_0 = 1 \) and \( n_e = n_{em} \). The conditions for the existence of solitary wave solutions are [20, 21]:

(i) \( V(n_0) = 0 \) at \( n_0 = 1 \) and \( n_e = n_{em} \)

(ii) \( \frac{dV}{dn_e} \bigg|_{n_{em}} = \frac{d^2V}{dn^2_e} \bigg|_{n_{em}} < 0 \)

(iii) \( V(n_e) < 0 \) for \( 1 < n_e < n_{em} \) (hump soliton)

For \( n_{em} < n_e < 1 \) dip soliton is formed. The particle is reflected back at \( n_0 = n_{em} \) so that \( V(n_e) > 0 \) for \( n_e > n_{em} \) (or \( n_e < n_{em} \) if \( n_{em} > 1 \) (or \( n_{em} < 1 \)).

As the nature of the Sagdeev potential \( V(n_e) \) determines the conditions for existence and nature of the solitary wave, it is important to study the nature of variation of \( V(n_e) \) for different plasma parameters such as positron concentration, plasma \( \beta \) and obliqueness of wave propagation. For numerical analysis we have chosen parameter ranges keeping in mind experimental as well as space and astrophysical situations.

To show the region of existence of solitary wave solution \( V(n_e) \) is drawn against \( n_e \) for different values of positron concentration (Fig. 1). It indicates hump type soliton solution whose amplitude decreases with increase in positron concentration. For \( p < 0.070747 \), the soliton solution does not exist.

In Fig. 2 we show the plot of \( V(n_e) \) versus \( n_e \) for different values of \( \beta \). It again shows excitation of hump type soliton. Amplitude of this soliton increases with increase in \( \beta \). However there is an upper limit of \( \beta \) (=0.150043) beyond which hump solitons cannot exist.

Fig. 1 Sagdeev profile for different values of positron concentration. The curves labeled \( a, b, c \) and \( d \) correspond respectively to \( p = 0.070647, \ 0.070747, \ 0.070847 \) and \( 0.070947 \)

Fig. 2 Sagdeev profile for different values of plasma \( \beta \). The curves labeled \( a, b, c \) and \( d \) correspond respectively to \( \beta = 0.149963, \ 0.150003, \ 0.150043 \) and \( 0.15083 \)

Fig. 3 Sagdeev profile for various values of the obliquity factor \( k_\parallel \). The curves labeled \( a, b, c \) and \( d \) correspond respectively to \( k_\parallel = 0.3594635, \ 0.3595335, \ 0.3596135 \) and \( 0.3596935 \)
Fig. 3 shows the plot of $V(n_{a})$ for various values of the obliquity factor $k_{a}$. Obviously the amplitude of these solitons depends sensibly on the propagation angle with respect to the external magnetic field. However, there is a lower limit to the value of $k_{a}$ (i.e., upper limit to the angle of propagation) beyond which hump solitons cannot exist. To summarize we have investigated in detail nonlinear slow shear Alfvén wave in electron-positron-ion plasma including ion motion parallel to the external magnetic field. We have examined the effects of positron concentration, strength of magnetic field and propagation angle with respect to the external magnetic field on the conditions of existence and nature of the slow shear Alfvén waves in electron-positron-ion plasmas. We have also examined the parametric regions of existence of nonlinear slow shear Alfvén solitary waves. There exists a lower limit of positron concentration ($\rho = 0.070747$) below which soliton solution does not exist. There exists an upper limit to the value of $\beta$ ($=0.150043$) beyond which hump solitons cannot exist. It has also been shown that the amplitude of slow shear Alfvén wave soliton depends sensibly on the propagation angle with respect to the external magnetic field and there is a lower limit to the value of $k_{a}$ (i.e., upper limit to the angle of propagation) beyond which hump solitons cannot exist. Earlier authors [16] studied the same problem ignoring ion parallel motion and predicted dip type soliton. Here we find that inclusion of ion parallel motion plays an important role and changes the nature of soliton from dip type to hump type. So, one must include full ion dynamics in order to study nonlinear slow shear Alfvén wave.

Finally, we would like to point out that the present investigation may be useful in many astrophysical and space environments where slow shear Alfvén wave may be excited in electron-positron-ion plasmas. Our investigation has been restricted to cold, homogeneous and collisionless plasma. The analysis can be generalized to include the effects of temperature, inhomogeneity and collision which are beyond the scope of the present work.

REFERENCES


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